# Trajectory Design Employing Convex Optimization for Landing on Irregularly Shaped Asteroids

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#### Goal

- Goal: Design an optimal powered descent trajectory on-board the spacecraft in order to softly land on an irregularly shaped asteroid.
  - Algorithm needs to be autonomous, reliable, robust, and efficient.
  - Designing on-board facilitates an easy change of parameters.
- Convex optimization is efficient and reliable.
  - Guarantees global minimum in a finite number of steps, if the problem is feasible.
  - Subclasses include Second Order Cone Programming (SOCP).
- Can convex optimization be used to design the asteroid powered descent trajectory?

# Original Problem Formulation

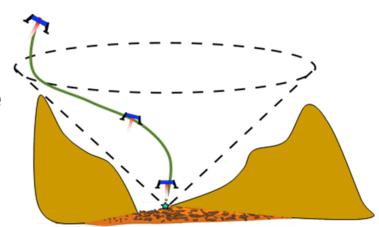
 Asteroid powered descent propellant optimal problem is nonlinear and nonconvex.

$$s.t. \ \dot{\vec{r}} = \vec{v}, \ \dot{\vec{v}} = \boxed{\vec{T}_{m}} - 2\vec{\omega} \times \vec{v} - \dot{\vec{\omega}} \times \vec{r} - \vec{\omega} \times (\vec{\omega} \times \vec{r}) + \boxed{\nabla U(\vec{r})}, \ \dot{m} = -\frac{1}{v_{ex}} \boxed{\parallel \vec{T} \parallel}$$

$$T_{min} \leq \boxed{\parallel \vec{T} \parallel} \leq T_{max}, \ \lVert \vec{r} - \vec{r}_{f} \rVert \cos\theta - (\vec{r} - \vec{r}_{f})^{T} \hat{n} \leq 0, \ m \geq m_{dry}$$

$$\vec{r}(0) = \vec{r}_{0}, \ \vec{v}(0) = \vec{v}_{0}, \ m(0) = m_{wet}, \ \vec{r}(t_{f}) = \vec{r}_{f}, \ \vec{v}(t_{f}) = \vec{v}_{f}, \ t_{f} \ \text{given}$$

- Fixed final time two point value boundary problem
- State:  $\vec{r}$ ,  $\vec{v}$ , m
- Control:  $\vec{T}$
- Highlighted terms are not permissible for a convex optimization problem.



#### **Problem Relaxation**

#### Relax the problem by introducing a slack variable, $T_m$ .

Original Problem 
$$\min_{min} -m(t_f)$$

$$s.t. \ \dot{\vec{r}} = \vec{v}$$

$$\dot{\vec{v}} = \frac{\vec{T}}{m} - 2\vec{\omega} \times \vec{v} - \dot{\vec{\omega}} \times \vec{r}$$

$$- \vec{\omega} \times (\vec{\omega} \times \vec{r}) + \nabla U(\vec{r})$$

$$\dot{m} = -\frac{1}{v_{ex}} \|\vec{T}\|$$

$$T_{min} \le \|\vec{T}\| \le T_{max}$$

$$\|\vec{r} - \vec{r}_f\| \cos \theta - (\vec{r} - \vec{r}_f)^T \hat{n} \le 0$$

$$m \ge m_{dry}$$

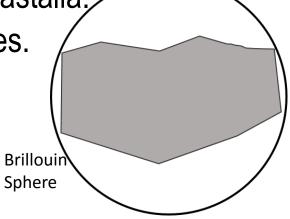
 $\vec{r}(0) = \vec{r}_0, \ \vec{v}(0) = \vec{v}_0, \ m(0) = m_{wet}$ 

 $\vec{r}(t_f) = \vec{r}_f, \ \vec{v}(t_f) = \vec{v}_f, \ t_f \text{ given}$ 

Proved the optimal solution of the relaxed problem is the optimal solution of the original.

# Irregularly Shaped Asteroid Gravity Models

- 4x4 Spherical Harmonics Model
  - Maximum Order and Degree 4
  - No symmetry nor coordinate system location and alignment assumptions.
  - High accuracy outside the Brillouin sphere.
  - Not valid inside the Brillouin sphere.
- Interior spherical Bessel gravity model
  - Valid inside the entire Brillouin sphere.
  - Error less than 10% for the binary asteroid Castalia,
  - Published in 2014 by Takahashi and Scheeres.



#### 4x4 Bessel

- 4x4 spherical harmonics gravity model outside the Brillouin sphere.
- Interior spherical Bessel gravity model inside the Brillouin sphere.
- Both models are summation series.
- Highly nonlinear in terms of spacecraft position vector.
- Computational similarities between the models allows for easy transition between the models.

#### Successive Solution Method

- Solve a series of convex optimization problems.
- Equations of motion can be arranged as:

$$\dot{\vec{x}}^{(k)} = A(\vec{r}^{(k-1)})\vec{x}^{(k)} + B\vec{u}^{(k)} + c(\vec{r}^{(k-1)})$$

- A and c are evaluated using the previous solution (k-1).
- In the (k)<sup>th</sup> iteration, dynamics are linear and time varying.
- Iterations continue until two successive trajectories are within a set tolerance.
- This is not the same as conventional linearization, as there are no approximations in the final iteration.
- Dominant gravity term is placed in A, with the higher order gravity terms in c.

#### Successive Solution Method: A, B, c

$$c = \begin{bmatrix} 0 \\ 0 \\ 0 \\ \frac{\partial U}{\partial r_x} - dom \\ \frac{\partial U}{\partial r_y} - dom \\ \frac{\partial U}{\partial r_z} - dom \\ 0 \end{bmatrix}^{(k-1)}$$
 Formulation assumes rotation vector is along the +Z axis,  $\vec{\omega} = \omega \hat{z}$ .

$$dom = \begin{cases} -\frac{\mu}{r^3} & 4 \times 4\\ \alpha_{0,0} j_1 \left(\frac{\alpha_{0,0} r}{R_b}\right) \bar{A}_{0,0,0} + \alpha_{1,0} j_1 \left(\frac{\alpha_{1,0} r}{R_b}\right) \bar{A}_{1,0,0} + \alpha_{2,0} j_1 \left(\frac{\alpha_{2,0} r}{R_b}\right) \bar{A}_{2,0,0} & Bessel \end{cases}$$

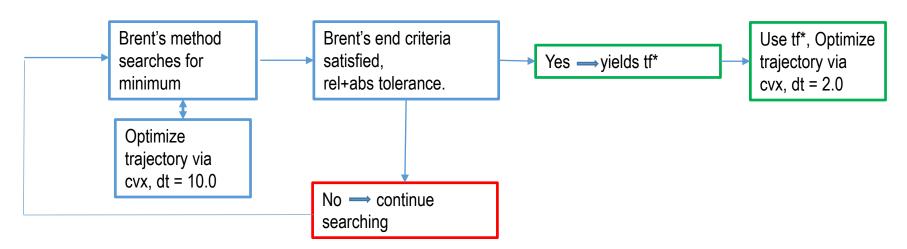
#### Additional Techniques

- Change of Variables
- Discretization
  - Continuous equation of motion turned into discrete equality constraints.
- Scaling

- Final optimization problem is convex.
  - Linear equality constraints
  - Convex inequality constraints
  - Inequality constraints are second order cone.
  - Actually a SOCP

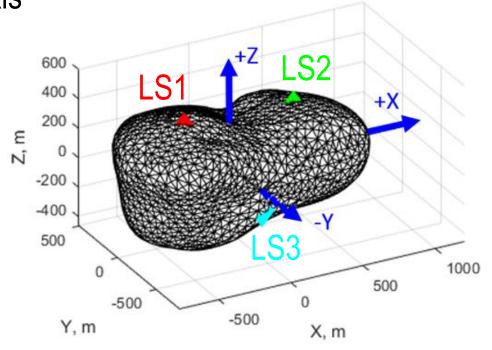
# **Optimal Flight Time**

- Desire to find the optimal flight time corresponding to smallest propellant usage.
- Propellant usage is unimodal with respect to flight time.
- Create an outer optimization loop using Brent's method to optimize the flight time.
- Use dt = 10.0 sec to find the optimal flight time. Design the final trajectory with dt = 2.0 sec.

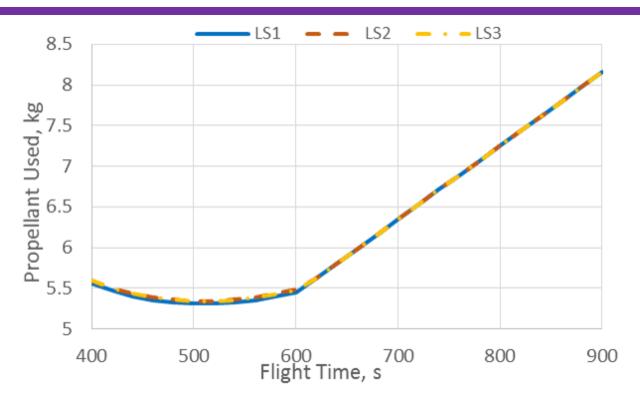


#### Simulation Parameters

- Asteroid Castalia
  - Period 4.07 hr along +Z axis
  - Three Landing Sites
- Spacecraft:
  - Mass 1400 kg
     ♦ 400 kg propellant
  - Thrust 80 N 20 N
- Initial Conditions
  - 1000 m altitude
  - Out of plane and uprange position and velocity components.
- Using CVX, a publically available Matlab based convex optimization program.



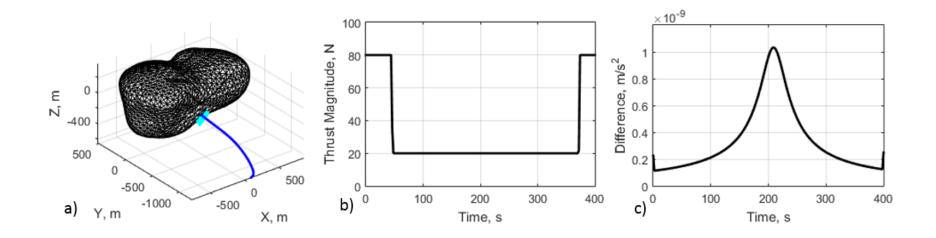
# Flight Time Parameter Sweep



- Typically 3 iterations required in the successive solution method.
  - Range 3 − 7
- Low number of iterations demonstrates stability in the successive solution methodology.

# Inner Loop Trajectory Design

- 400 Second Flight Time landing at LS3
- Thrust profile follows the traditional bang-bang.

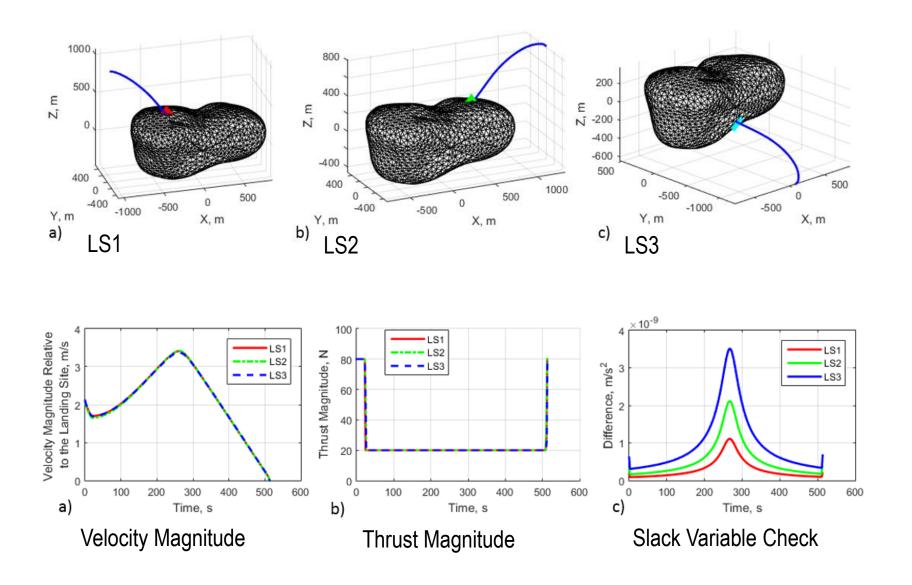


#### Optimal Flight Time Optimal Propellant Trajectory

Combined outer and inner loop executions took
 2.2 - 2.5 minutes.

	Optimal Flight Time, s	Propellant Used, kg	Number of Inner Loop Executions
LS1	512.86	5.31	7
LS2	512.27	5.34	7
LS3	513.35	5.34	7

# Optimal Flight Time Optimal Propellant Trajectory Parameters

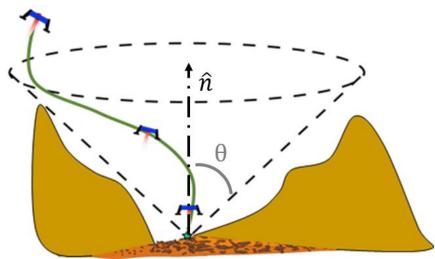


# Glide Slope Constraint

 Glide slope constraint: Constrains the vehicle to fly inside a cone around the landing site.

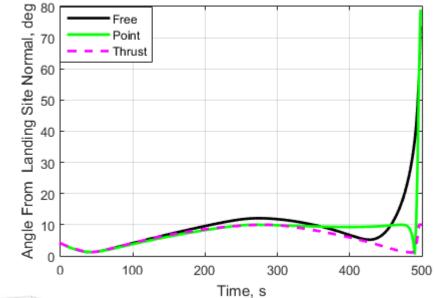
$$\bullet \|\vec{r} - \vec{r}_f\| \cos \theta - (\vec{r} - \vec{r}_f)^T \hat{n} \le 0$$

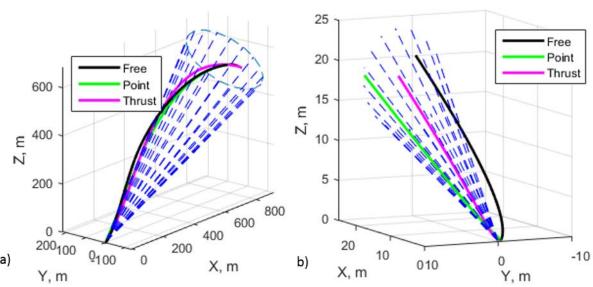
- Near the landing site the vehicle must match the landing site velocity to rotate with the landing site.
- Low thrust of the vehicle (80 N) prohibits this.
- Alternate solutions for a 10 deg cone:
  - Increase max thrust to 320 N
  - Enforce the constraint for all, but the last 6 seconds.
    - ♦ Flies slightly outside the cone near the surface.



# Glide Slope Results

- LS2 500 second flight time.
- 10 degree cone enforced.





#### Conclusions

- Asteroid powered descent trajectory design can be formulated as a convex optimization problem.
- Successive solution methodology is the key to handling a nonlinear gravity model.
- Formulated algorithm handles a wide range of parameters successfully.
- Flight time optimization is completed in an outer loop with Brent's method.
- Inclusion of additional trajectory constraints in the algorithm is feasible.
- Viable algorithm for rapidly designing asteroid powered descent trajectories autonomously on-board the spacecraft for use in a variety of guidance algorithms.

# **BACK-UP**

# Convex Optimization and SOCP Formulation

• Optimization problem formulation min g(x)

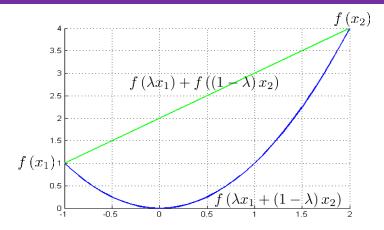
s.t. 
$$f_i(x) \le 0$$
  $i = 1, ..., m$   
 $h_j(x) = 0$   $j = 1, ..., p$ 

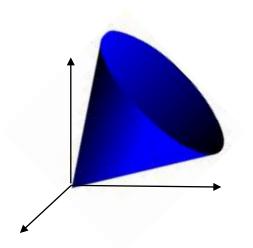


- g(x) and  $f_i(x)$  are convex functions.
- $h_j(x)$  is linear.
- Convex Function:  $f(\lambda x_1 + (1 \lambda) x_2) \le f(\lambda x_1) + f((1 \lambda) x_2)$

#### Second Order Cone Program (SOCP)

- Subset of convex optimization
- g(x) and  $h_i(x)$  are linear functions.
- $f_i(x)$  is second order cone.
- Second order Cone:  $||Mx + d||_2 \le c$





# Spherical Harmonics Gravity Model

Fidelity determined by the coefficients and the number of terms in the summation series.

$$N = \begin{cases} 2 & 2 \times 2 \\ 4 & 4 \times 4 \end{cases}$$

Cartesian coordinate system:

$$\frac{\partial U}{\partial r_x} = \frac{\partial U}{\partial r} \frac{\partial r}{\partial x} + \frac{\partial U}{\partial \delta} \frac{\partial \delta}{\partial x} + \frac{\partial U}{\partial \lambda} \frac{\partial \lambda}{\partial x}$$

$$\frac{\partial U}{\partial r_y} = \frac{\partial U}{\partial r} \frac{\partial r}{\partial y} + \frac{\partial U}{\partial \delta} \frac{\partial \delta}{\partial y} + \frac{\partial U}{\partial \lambda} \frac{\partial \lambda}{\partial y}$$

$$\frac{\partial U}{\partial r_z} = \frac{\partial U}{\partial r} \frac{\partial r}{\partial z} + \frac{\partial U}{\partial \delta} \frac{\partial \delta}{\partial z} + \frac{\partial U}{\partial \lambda} \frac{\partial \lambda}{\partial z}$$

 $P_{l,m}$  associated Legendre function l. m order, degree  $r, \delta, \lambda$  radius, latitude, longitude

Partial of the gravitational potential with respect to the position vector in spherical coordinates:

$$N = \begin{cases} 2 & 2 \times 2 \\ 4 & 4 \times 4 \end{cases}$$

$$\frac{\partial U}{\partial r} = \sum_{l=0}^{N} \sum_{m=0}^{l} -(l+1) \frac{\mu}{r^2} \left(\frac{r_0}{r}\right)^l P_{l,m} \left[\sin \delta\right] \left\{C_{l,m} \cos \left(m\lambda\right) + S_{l,m} \sin \left(m\lambda\right)\right\}$$

$$\frac{\partial U}{\partial r} = \sum_{l=0}^{N} \sum_{m=0}^{l} -(l+1) \frac{\mu}{r^2} \left(\frac{r_0}{r}\right)^l P_{l,m} \left[\sin \delta\right] \left\{C_{l,m} \cos \left(m\lambda\right) + S_{l,m} \sin \left(m\lambda\right)\right\}$$

$$\frac{\partial U}{\partial r} = \sum_{l=0}^{N} \sum_{m=0}^{l} \frac{\mu}{r} \left(\frac{r_0}{r}\right)^l \left\{C_{l,m} \cos \left(m\lambda\right) + S_{l,m} \sin \left(m\lambda\right)\right\} \frac{\partial P_{l,m} \left[\sin \delta\right]}{\partial \delta}$$
Cartesian coordinate
$$\frac{\partial U}{\partial \lambda} = \sum_{l=0}^{N} \sum_{m=0}^{l} \frac{\mu}{r} \left(\frac{r_0}{r}\right)^l m P_{l,m} \left[\sin \delta\right] \left\{-C_{l,m} \sin \left(m\lambda\right) + S_{l,m} \cos \left(m\lambda\right)\right\}$$

Partial of the position vector in spherical coordinate system with respect to the Cartesian:

$$\frac{\partial r}{\partial x} = \frac{r_x}{r}, \ \frac{\partial r}{\partial y} = \frac{r_y}{r}, \ \frac{\partial r}{\partial z} = \frac{r_z}{r}$$

$$\frac{\partial \delta}{\partial x} = \frac{-r_x r_z}{r^2 \sqrt{r_x^2 + r_y^2}}, \ \frac{\partial \delta}{\partial y} = \frac{-r_y r_z}{r^2 \sqrt{r_x^2 + r_y^2}}, \frac{\partial \delta}{\partial z} = \frac{1}{\sqrt{r_x^2 + r_y^2}} \left(1 - \frac{r_z^2}{r^2}\right)$$

$$\frac{\partial \lambda}{\partial x} = \frac{-r_y}{r_x^2 + r_y^2}, \ \frac{\partial \lambda}{\partial x} = \frac{r_x}{r_x^2 + r_y^2}, \ \frac{\partial \lambda}{\partial z} = 0$$

# Interior spherical Bessel Gravity Model

■ Summation Series:  $l_{max} = 2$ ,  $n_{max} = 5$ ,  $m_{max} = 5$ 

$$\frac{\partial U}{\partial r_{x}} = \frac{\mu}{R_{b}} \sum_{l=0}^{l_{max}} \sum_{n=0}^{n_{max}} \sum_{m=0}^{n} Re \left[ \frac{\partial}{\partial x} \left( \bar{\beta}_{n,m} \left( \alpha_{l,n} \right) \right) \right] \bar{A}_{l,n,m} + Im \left[ \frac{\partial}{\partial x} \left( \bar{\beta}_{n,m} \left( \alpha_{l,n} \right) \right) \right] \bar{B}_{l,n,m}$$

$$\frac{\partial U}{\partial r_{y}} = \frac{\mu}{R_{b}} \sum_{l=0}^{l_{max}} \sum_{n=0}^{n_{max}} \sum_{m=0}^{n} Re \left[ \frac{\partial}{\partial y} \left( \bar{\beta}_{n,m} \left( \alpha_{l,n} \right) \right) \right] \bar{A}_{l,n,m} + Im \left[ \frac{\partial}{\partial y} \left( \bar{\beta}_{n,m} \left( \alpha_{l,n} \right) \right) \right] \bar{B}_{l,n,m}$$

$$\frac{\partial U}{\partial r_{z}} = \frac{\mu}{R_{b}} \sum_{l=0}^{l_{max}} \sum_{n=0}^{n_{max}} \sum_{m=0}^{n} Re \left[ \frac{\partial}{\partial z} \left( \bar{\beta}_{n,m} \left( \alpha_{l,n} \right) \right) \right] \bar{A}_{l,n,m} + Im \left[ \frac{\partial}{\partial z} \left( \bar{\beta}_{n,m} \left( \alpha_{l,n} \right) \right) \right] \bar{B}_{l,n,m}$$

■ Basis Functions:  $\bar{\beta}_{n,m}(\alpha_{l,n}) = j_n \left[ \frac{\alpha_{l,n} r}{R_b} \right] \bar{H}_{n,m}$ 

$$\bar{H}_{n,m} = \begin{cases} N_{n,m} P_{n,m} \left[ sin(\phi) \right] e^{im\lambda} & n \ge m \ge 0 \\ 0 & otherwise \end{cases}$$

Partials of the Basis Functions:

$$\begin{split} \frac{\partial}{\partial x} \left( \bar{\beta}_{n,m} \left( \alpha_{l,n} \right) \right) &= \begin{cases} -\frac{\alpha_{l,n} r_x}{R_b r} j_{n+1} \begin{bmatrix} \alpha_{l,n} r \\ R_b \end{bmatrix} \bar{H}_{n,0} - 2 \bar{\mathscr{F}}_1 \left( n, m \right) \frac{1}{r} j_n \begin{bmatrix} \alpha_{l,n} r \\ R_b \end{bmatrix} N_{n-1,1} P_{n-1,1} \left[ \sin(\phi) \right] \cos \lambda & m = 0 \\ -\frac{\alpha_{l,n} r_x}{R_b r} j_{n+1} \begin{bmatrix} \alpha_{l,n} r \\ R_b \end{bmatrix} \bar{H}_{n,m} - \bar{\mathscr{F}}_1 \left( n, m \right) \frac{1}{r} j_n \begin{bmatrix} \alpha_{l,n} r \\ R_b \end{bmatrix} \bar{H}_{n-1,m+1} + \bar{\mathscr{F}}_2 \left( n, m \right) \frac{1}{r} j_n \begin{bmatrix} \alpha_{l,n} r \\ R_b \end{bmatrix} \bar{H}_{n-1,m-1} & m > 0 \\ \frac{\partial}{\partial y} \left( \bar{\beta}_{n,m} \left( \alpha_{l,n} \right) \right) &= \begin{cases} -\frac{\alpha_{l,n} r_y}{R_b r} j_{n+1} \begin{bmatrix} \alpha_{l,n} r \\ R_b \end{bmatrix} \bar{H}_{n,0} - 2 \bar{\mathscr{F}}_1 \left( n, m \right) \frac{1}{r} j_n \begin{bmatrix} \alpha_{l,n} r \\ R_b \end{bmatrix} N_{n-1,1} P_{n-1,1} \left[ \sin(\phi) \right] \sin \lambda & m = 0 \\ -\frac{\alpha_{l,n} r_y}{R_b r} j_{n+1} \begin{bmatrix} \alpha_{l,n} r \\ R_b \end{bmatrix} \bar{H}_{n,m} + \bar{\mathscr{F}}_1 \left( n, m \right) \frac{1}{r} j_n \begin{bmatrix} \alpha_{l,n} r \\ R_b \end{bmatrix} i \bar{H}_{n-1,m+1} + \bar{\mathscr{F}}_2 \left( n, m \right) \frac{1}{r} j_n \begin{bmatrix} \alpha_{l,n} r \\ R_b \end{bmatrix} i \bar{H}_{n-1,m-1} & m > 0 \\ \frac{\partial}{\partial z} \left( \bar{\beta}_{n,m} \left( \alpha_{l,n} \right) \right) &= -\frac{\alpha_{l,n} r_z}{R_b r} j_{n+1} \begin{bmatrix} \alpha_{l,n} r \\ R_b \end{bmatrix} \bar{H}_{n,m} + \bar{\mathscr{F}}_3 \left( n, m \right) \frac{1}{r} j_n \begin{bmatrix} \alpha_{l,n} r \\ R_b \end{bmatrix} \bar{H}_{n-1,m} \\ \frac{\partial}{\partial z} \left( \bar{\beta}_{n,m} \left( \alpha_{l,n} \right) \right) &= -\frac{\alpha_{l,n} r_z}{R_b r} j_{n+1} \begin{bmatrix} \alpha_{l,n} r \\ R_b \end{bmatrix} \bar{H}_{n,m} + \bar{\mathscr{F}}_3 \left( n, m \right) \frac{1}{r} j_n \begin{bmatrix} \alpha_{l,n} r \\ R_b \end{bmatrix} \bar{H}_{n-1,m} \end{aligned}$$

#### Thrust Profile

Three classes of thrust profiles

